

In this problem, we investigate the effect of electromagnetic waves traveling through a gas of charged particles. This can happen when there is radio emission from a pulsar, and these signals propagate through clouds of charged particles in deep space before being detected on Earth. A linearly polarized radio wave will induce a charged current in the cloud which is proportional to the time-dependent electric field of the plane wave (ignore the motion of the charged particles due to the magnetic field of the plane wave).

a. Show that the dispersion relation between the frequency ω and the wavevector k for plane waves traveling through an electron gas can be written in terms of

$$1 - \frac{\omega_p^2}{\omega^2}$$

where ω_p is the plasma frequency. Express the plasma frequency in terms of: $-e = -4.8 \times 10^{-10} \text{esu}$ (the electron charge) and n_e (the volume density of electrons in the cloud).

b. For radio wave frequencies above ω_p , how significant is the dispersion from ions (protons) in comparison to electrons?

c. Evaluate the phase velocity ω / k and the group velocity $d\omega/dk$ and compare them to the speed of light. Write the phase and group velocities in terms of the ratio ω / ω_p .

The Vela pulsar is about 500 parsecs distant (1 parsec = $3 \times 10^{18} \text{cm}$). It emits radio waves over a broad band. When observations are made in narrow frequency bands, what is observed are narrow pulses which arrive at a fixed period, similar to a timing signal for synchronizing a clock.

d. The narrow pulses observed at 1660 MHz are delayed relative to the narrow pulses observed at 1720 MHz by 6.8 ms. If this is interpreted by the dispersion in an ionized gas, what is the mean density of free electrons between Vela and us? To simplify the calculation, you can anticipate that $\omega_p \ll \omega$.

We start off with some of Maxwell's equations:

$$\nabla \times \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t} \quad \nabla \times \mathbf{B} = \frac{4\pi}{c} \mathbf{J} + \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t}$$

And the vector relation:

$$\nabla \times (\nabla \times \mathbf{A}) = \nabla (\nabla \cdot \mathbf{A}) - \nabla^2 \mathbf{A}$$

Taking the cross product of the first relation and using the vector identity:

$$\nabla \times (\nabla \times \mathbf{E}) = \nabla (\nabla \cdot \mathbf{E}) - \nabla^2 \mathbf{E} = -\frac{1}{c^2} \frac{\partial}{\partial t} \left(4\pi \mathbf{J} + \frac{\partial \mathbf{E}}{\partial t} \right)$$

We can rewrite the current as:

$$\mathbf{J} = n_e e \mathbf{v}$$

And writing the Lorentz force equation

$$m_e \mathbf{a} = -q \left(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right)$$

Since v_1 is a perturbed quantity and $v_0 = 0$, we can assume it is small compared to c and neglect the magnetic contribution to the Lorentz force:

$$\frac{\partial \mathbf{J}}{\partial t} = n_e e \frac{e}{m_e} \mathbf{E}$$

so that:

$$\nabla (\nabla \cdot \mathbf{E}) - \nabla^2 \mathbf{E} = -\frac{1}{c^2} n_e e \frac{e}{m_e} \mathbf{E} - \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2}$$

We can write the perturbed values in terms of plane waves moving in the \hat{z} direction, for example:

$$\mathbf{E} = \mathbf{E}_1 \exp i(kz - \omega t)$$

Here we have taken $E_0 = 0$ (no zero order electric field). Plugging in:

$$\nabla (ik\hat{z} \cdot \mathbf{E}) + k^2 \mathbf{E}_1 = -\frac{1}{c^2} \frac{n_e e^2}{m_e} \mathbf{E}_1 + \frac{1}{c^2} \omega^2 \mathbf{E}_1$$

The first term on the left must be zero, since the electric field must be perpendicular to the wave vector:

$$c^2 k^2 = \omega^2 - \frac{n_e e^2}{m_e}$$

Defining $\omega_p^2 = n_e e^2 / m_e$, we get the dispersion relation:

$$\omega^2 = \frac{c^2 k^2}{1 - \omega_p^2 / \omega^2}$$

For frequencies above ω_p , dispersion from ions have negligible dispersion compared to electrons, since the response

from ions will be like $\omega_{pi} = \omega_p \frac{m_e}{m_i} \sim \omega_p / 1800$ because of the mass in the frequency. This means that

$$\omega_{pe}^2 / \omega^2 \gg \omega_{pi}^2 / \omega^2.$$

The phase velocity is just:

$$\frac{\omega}{k} = \frac{1}{\sqrt{1 - \omega_p^2 / \omega^2}} = \frac{1}{\sqrt{1 - \omega_p^2 / (\omega_p^2 + k^2)}}$$

where for the second equality we use the dispersion relation again. The group velocity can be found by differentiating:

$$\omega^2 = k^2 + \omega_p^2 \Rightarrow 2\omega \frac{\partial \omega}{\partial k} = 2k$$

Moving around and substituting in the phase velocity:

$$\frac{\partial \omega}{\partial k} = \frac{k}{\omega} = \sqrt{1 - \omega_p^2 / \omega^2} = \sqrt{1 - \omega_p^2 / (\omega_p^2 + k^2)}$$

From part c:

$$v_g = \frac{\omega}{k} = \frac{1}{\sqrt{1 - \omega_p^2 / \omega^2}}$$

The time it takes a pulse of frequency ω to cross to here:

$$t(\omega) = \frac{d}{c} \sqrt{1 - \omega_p^2 / \omega^2}$$

Simplifying using $\omega_p / \omega \ll 1$:

$$t(\omega) = \frac{d}{c} \left(1 - \frac{\omega_p^2}{2\omega^2} \right)$$

So that the difference in time is given by:

$$\Delta t = \frac{d\omega_p^2}{2c} \left(\frac{1}{(1660\text{MHz})^2} - \frac{1}{(1720\text{MHz})^2} \right)$$

Plugging in for everything:

$$6.8 \times 10^{-3} \text{ s} \text{ amp;} = \text{amp;} \frac{(500 \times 3 \times 10^{18} \text{ cm}) n_e (4.8 \times 10^{-10} \text{ esu})^2}{2(3 \times 10^{10} \text{ cm/s})(9.1 \times 10^{-28} \text{ g})} \text{ amp;} \cdot \left(\frac{1}{(1.66 \times 10^9 \text{ s}^{-1})^2} - \frac{1}{(1.72 \times 10^9 \text{ s}^{-1})^2} \right)$$

and solving:

$$n_e = 4.3 \times 10^4 / \text{cm}^3$$